# Harmonic Oscillator

Chem 2340

## **Classical harmonic oscillator**

$$F_{x} = -kx$$

$$m \frac{d^{2}x}{dt^{2}} + kx = 0$$

$$x = A \sin(2\pi \upsilon t + b)$$

$$\upsilon = \frac{1}{2\pi} \sqrt{\frac{k}{m}}; \text{ or } \omega = 2\pi \upsilon = \sqrt{\frac{k}{m}} = \text{angular frequency}$$

$$V = -\int F_{x} dx = \int kx dx = \frac{1}{2} kx^{2}$$

$$KE = \frac{1}{2}m\left(\frac{dx}{dt}\right)^2$$
$$E = T + V = 2\pi^2\upsilon^2 mA^2 = \frac{1}{2}kA^2$$

$$k = 4\pi^2 \upsilon^2 m$$

Note: here we used  $\cos^2(y) + \sin^2(y) = 1$ 

#### Series solutions to DE's

Consider

 $y''+c^2y=0$  here  $y''=d^2y/dx^2$ 

by inspection, the solutions are

 $y_1 = \sin(cx), y_2 = \cos(cx)$ 

But we are not always going to be able to "guess" solutions.

A more general approach is to write

$$y = a_0 + a_1 x + a_2 x^2 + \dots = \sum_{n=0}^{\infty} a_n x^n$$

So the DE becomes.

$$\sum_{n=2}^{\infty} ann(n-1)x^{n-2} + c^2 \sum_{n=0}^{\infty} a_n x^n = 0$$

This can be rewritten as

$$\sum_{n=0}^{\infty} (n+2)(n+1)a_{n+2}x^n + c^2 \sum_{n=0}^{\infty} a_n x^n = 0$$

$$y' = \sum_{n=1}^{\infty} na_n x^{n-1}$$

$$y''=\sum_{n=2}^{\infty}n(n-1)x^{n-2}$$

$$\sum_{n=0} \left[ (n+2)(n+1)a_{n+2} + a_n c^2 \right] x^n = 0$$

$$\Rightarrow (n+2)(n+1)a_{n+2} + c^2a_n = 0$$

$$2 \cdot 1a_2 + c^2 a_0 = 0$$
$$3 \cdot 2a_3 + c^2 a_1 = 0$$
$$4 \cdot 3a_4 + c^2 a_2 = 0$$

So  $a_2$ ,  $a_4$ ,  $a_6$ ,... are in terms of  $a_o$  $a_3$ ,  $a_5$ ,  $a_7$ ,... are in terms of  $a_1$ 

What is the significance of this?

Note: the equation relating higher  $n a_n$  to the lower n coefficients is a recursion relation.

Now let's consider the SE for the harmonic oscillator

$$\frac{-\hbar^2}{2m}\frac{d^2}{dx^2}\psi + V\psi = E\psi$$

$$\frac{d^2\psi}{dx^2} + \left(\frac{2mE}{\hbar^2} - \alpha^2 x^2\right)\psi = 0$$

Actually, for a diatomic molecule the mass used would be the reduced mass.

$$\mu = \frac{m_1 m_2}{m_1 + m_2}$$

$$V = 1/2kx^{2}$$
$$\upsilon = \frac{1}{2\pi}\sqrt{\frac{k}{m}};$$
$$4\pi^{2}\upsilon^{2}m = k$$
$$\frac{4\pi^{2}\upsilon^{2}m^{2}}{\hbar^{2}} = \alpha^{2}$$

$$\alpha = 2\pi \upsilon m / \hbar$$

$$\frac{d^2\psi}{dx^2} + \left(\frac{2mE}{\hbar^2} - \alpha^2 x^2\right)\psi = 0$$

Could directly solve using the series method Alternatively, we can show that  $e^{-\alpha x^2/2}$  is a solution

$$\psi = e^{-\alpha x^{2}/2}$$

$$\psi' = -\alpha x e^{-\alpha x^{2}/2}$$

$$\psi'' = \left(-\alpha + \alpha^{2} x^{2}\right) e^{-\alpha x^{2}/2}$$

$$\left[-\alpha + \alpha^{2} x^{2} + \frac{2mE}{\hbar^{2}} - \alpha^{2} x^{2}\right] e^{-\alpha x^{2}/2} = 0 \implies$$

$$\alpha = \frac{2mE}{\hbar^{2}}, E = \frac{2\pi \nu m}{\hbar} \frac{\hbar^{2}}{2m} = \pi \nu \hbar = \frac{\hbar \omega}{2} = \frac{1}{2} h\nu$$

We then look for a general solution of the

$$\psi(x) = e^{-\alpha x^2/2} f(x)$$

We discard the  $e^{\alpha x^2/2}$  solution

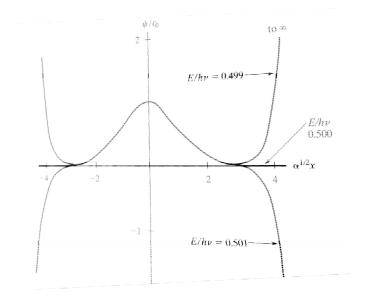


Fig 4.2 from text showing impact of small error in the energy on the wave function In a numerical solution to the problem

$$f''-2\alpha x f' + \left(\frac{2mE}{\hbar^2} - \alpha\right) f = 0$$
  
Let  $f = \sum_{n=0}^{\infty} c_n x^n$   
can show  $c_{n+2} = \frac{\alpha + 2\alpha n - 2mE/\hbar^2}{(n+1)(n+2)} c_n$ 

Boundary conditions are  $\psi \rightarrow 0$  as  $x \rightarrow \pm \infty$ 

So f(x) must not grow faster than  $e^{-\alpha x^2/2}$  falls of f as  $x \to \pm \infty$ 

Ratio test 
$$\frac{c_{2n+2}}{c_n} \sim \frac{\alpha}{n}$$
, for large *n*

Can show this implies that the series behaves as  $e^{\alpha x^2}$ 

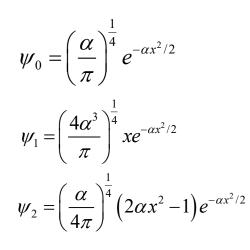
So we have a problem. We can solve this by requiring that f(x) is a finite polynomial rather than an  $\infty$  series.

For the series to truncate some n = v we require

$$\alpha + 2\alpha v - 2mE / \hbar^{2} = 0$$
  
Then  $E = \frac{(\alpha + 2\alpha v)}{2m} \hbar^{2} = \frac{2\alpha \left[\frac{1}{2} + v\right]}{2m} \hbar^{2}$ 
$$= \frac{2(2\pi v)}{2m\hbar} \left[\frac{1}{2} + v\right] \hbar^{2} = 2\pi v \hbar \left[\frac{1}{2} + v\right] = \hbar \omega \left[\frac{1}{2} + v\right]$$

or 
$$\hbar\omega \left[ n + \frac{1}{2} \right]$$
 If we switch back  
to the *n* variable  
So the energies are  $\frac{1}{2}\hbar\omega, \frac{3}{2}\hbar\omega, \frac{5}{2}\hbar\omega, \dots$ 

$$c_{n+2} = \frac{2\alpha (n-\nu)}{(n+1)(n+2)}c_n$$



The polynomials in this sequence are the Hermite polynomials,  $\{H_i\}$ 

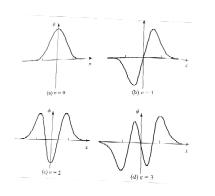


Fig. 4.4 from text showing the wave functions of the first four levels of the HO

$$\langle E \rangle_i = \langle \psi_i | H | \psi_i \rangle = \langle T \rangle_i + \langle V \rangle_i$$

Just as for the classical HO  $\langle T \rangle = \langle V \rangle$ 

Note it is the  $\{\psi_i\}$  that form an orthogonal set, not the  $H_i$   $\langle \psi_i | \psi_j \rangle = \int_{-\infty}^{\infty} H_i H_j e^{-\alpha x^2} dx = \delta_{ij}$   $\langle x \rangle_i = 0$  Do you see why?  $\langle x^2 \rangle_i \neq 0$   $\langle p_x \rangle_i = 0$  $\langle p_x^2 \rangle_i \neq 0$  There are lots of powerful relations involving Hermite polynomials.

Two are

 $H_{n}(z) = (-1)^{n} e^{z^{2}} \frac{d^{n}}{dz^{n}} e^{-z^{2}} \text{ generating function}$   $2zH_{n}(z) = nH_{n-1}(z) + \frac{1}{2}H_{n+1}(z) \text{ recursion relationship}$   $H_{0} = 1$   $H_{1} = 2z$   $H_{2} = 4z^{2} - 2$   $H_{3} = 8z^{3} - 12z$ 

 $z = \sqrt{\alpha}x$  to get the polynomials in terms of  $\alpha$  and x.

## Transition matrix element for going from level 0 to level n

The operator corresponding to radiation in the z direction is proportional to z

$$= \langle 0 | z | n \rangle = \int_{-\infty}^{\infty} \psi_0 z \psi_n dz = \int_{-\infty}^{\infty} H_0 \cdot z H_n(z) e^{-z^2} dz$$
  
$$= \int_{-\infty}^{\infty} \left[ \frac{1}{2} H_0 n H_{n-1}(z) + \frac{1}{4} H_0 H_{n+1}(z) \right] e^{-z^2} dz$$
  
$$= n \int_{-\infty}^{\infty} H_0 H_{n-1}(z) e^{-z^2} dz + \frac{1}{2} \int_{-\infty}^{\infty} H_0 H_{n+1}(z) e^{-z^2} dz$$
  
Non-zero only if  $n=1$  0

Here I have not normalized the wave functions. To get the correct value of the integral, when non-zero, one has to use normalized functions and to include the prefactor of the operator

If starting in level *n*, the allowed transitions are to  $n\pm 1$ 

Intensity  $\propto |\langle 0 | z | n \rangle|^2$ 

In the Dirac notation the identity operator I can be written as:

$$\sum_{n=0}^{\infty} \left| n > < n \right|$$

Let's check

$$\sum_{n=0} |n > < n| \sum_{m=0} |m > < m| = \sum_{n,m} |n > < n | m > < m|$$
$$= \sum_{n=0} |n > < n| \text{ since } \langle n | m \rangle = \delta_{nm}$$
Also  $\left( \sum_{n=0}^{\infty} |n > < n| \right) |m > = |m >$ 

Note that the identify operator involves an expansion over all the eigenfunctions of an operator

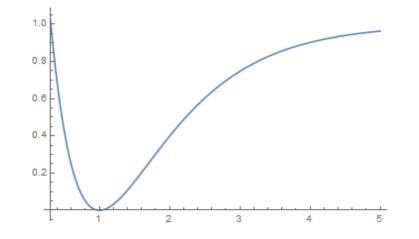
Supposed you want to evaluate:  $\langle 0 | z^2 | 0 \rangle$ 

We can write: 
$$\langle 0 | z^2 | 0 \rangle = \langle 0 | zIz | 0 \rangle = \sum_{n=0}^{\infty} \langle 0 | z | n \rangle \langle n | z | 0 \rangle$$
  
=  $|\langle 0 | z | 1 \rangle|^2$ 

 $\hbar\omega$ 

### Realistic potential energy curves of diatomic molecules

For a real molecule the potential energy curve bends over and  $V \rightarrow D_e$  as  $x \rightarrow \infty$ 

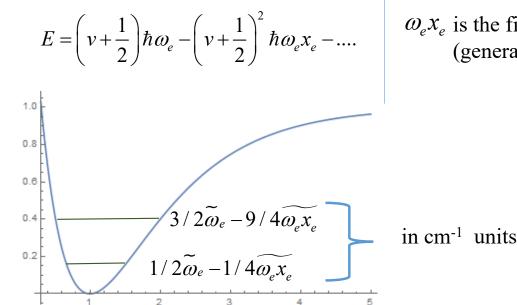


Taylor series expansion about its minimum at  $x = x_0$ 

$$V(x) = \frac{d^2 V / dx^2}{2} (x - x_e)^2 + \frac{d^3 V / dx^3}{3!} (x - x_e)^3 + \frac{d^4 V / dx^4}{4!} (x - x_e)^4 + \dots$$

Thus a harmonic potential represents the leading term of the Taylor series expansion of the potential about its minimum at  $x_{e}$ .

Can show that:



 $\omega_e x_e$  is the first unharmonicity constant (generally positive)

Population of levels at finite T

 $\frac{N_i}{N_i} = \frac{g_i}{g_i} e^{-(E_i - E_j)/kT}$ Boltzmann factor in exponent

where  $g_i$  is the degeneracy of level *i* which is 1 for the 1D vibrational problem.

For chemically bonded diatomic molecules, only the n=0 level has significant population at room temperature.

kT at room temperature ~200 cm<sup>-1</sup>. For typical diatomics,  $\hbar\omega$  (in cm<sup>-1</sup> units) is 2000-3000 cm<sup>-1</sup>.